



- A freely-falling Calvin

This week we have a little Christoffel computing, work with geodesics on a sphere, and more on accelerating observers. It looks like we are on track to assemble Einstein's equation before break. After break we will study the main current applications: Gravitational waves, black holes, and cosmology.

Reading:

- We discussed $T^{\mu\nu}$ in Schutz Chapter 4 pages 84-89, sections 4.2 and 4.6 and page 100. We skipped the rest.
- Schutz Chapter 5 sections 1 - 4 (We skip section 5.5.)
- A look ahead - Schutz Chapter 6 on Riemannian geometry

Problems: Solutions due by 11 PM on Thursday, March 5

All numbered problems are from Schutz.

- (1) "Oops!" What is wrong with the following equation?

$$\frac{dp^\mu}{d\tau} = qF^{\mu\nu}U^\beta$$

Using the metric $g_{\alpha\beta}$ how might you correct the typo?

- (2) GPS satellites emit signals at a constant rate, as measured by an onboard clock. They orbit at an altitude of 20.2×10^6 m.
- (a) Using the approximate relativistic correction you found in PS 1 problem 7 find the relative fractional difference

$$\frac{\Delta t - \Delta \bar{t}}{\Delta t}$$

in the clock rates onboard the GPS satellites and on Earth. GPS satellites travel at $v_s = \sqrt{GM}/r_s$, where r_s is the radius of the satellite's orbit. The speed of the surface of Earth in the (approximate) inertial frame at Earth's center is 0.46 km/s.

- (b) Using the effect of gravity on clock rates we derived in class, find the fractional difference $(\Delta\tau_B - \Delta\tau_A)/\Delta\tau_A$ due to the gravitational potential.
- (c) Find the net effect. Comment on the results.
- (3) 5.2 On tides
- (4) 5.7 Computing the transformation as we did in class

- (5) 5.8 Computing with 1-forms and vectors in 2D
- (6) DELAYED TO NEXT WEEK 5.21 (2 pts.) The last of the “Rindler spacetime set” of 2.19, 2.21, and 5.21.
- (7) Consider the 2D spacetime metric with coordinates (t, r) ,

$$ds^2 = - \left(1 - \frac{1}{\sqrt{r^2 + a^2}} \right) dt^2 + \left(1 - \frac{1}{\sqrt{r^2 + a^2}} \right)^{-1} dr^2,$$

where a is a constant.

- (a) Write the metric and its inverse as 2×2 arrays in these coordinates.
- (b) Compute the non-vanishing Christoffel symbols. Please show your work.
- (8) DELAYED TO NEXT WEEK Compute the Christoffel symbols Γ_{tt}^t and Γ_{tt}^i for the metric shown in equation (7.8) assuming that the function $\phi = \phi(t, x, y, z)$ is a function of all the coordinates.

